# Solvable symmetric black holes in anti-de Sitter spaces 

Laurent Claessens ${ }^{\text {a,*,1 }}$, Stéphane Detournay ${ }^{\text {b,1 }}$<br>${ }^{\text {a }}$ Département de Mathématique, Université Catholique de Louvain, Chemin du cyclotron, 2, 1348 Louvain-La-Neuve, Belgium<br>${ }^{\text {b }}$ Mécanique et Gravitation, Université de Mons-Hainaut, 20 Place du Parc, 7000 Mons, Belgium

Received 28 April 2006; received in revised form 2 July 2006; accepted 3 August 2006
Available online 1 September 2006
Dedicated to Jeanne


#### Abstract

Using symmetric space techniques, we show that closed orbits of Iwasawa subgroups of $S O(2, l-1)$ naturally define the singularity of a black hole causal structure in anti-de Sitter spaces in $l \geq 3$ dimensions. In particular, we recover for $l=3$ the non-rotating massive BTZ black hole. The method presented here is very simple and in principle generalizable to any semi-simple symmetric space. © 2006 Elsevier B.V. All rights reserved.


JGP SC: Real and complex differential geometry; Lie groups; General relativity
MSC: 53C35; 83C57
Keywords: Symmetric spaces; Anti-de Sitter spaces; Black holes; Minimal parabolic subgroups

## 1. Introduction

From a geometric point of view, a black hole is the data of a time orientable pseudo-Riemannian manifold $M$ together with a subset $\mathscr{S} \subset M$ called a singularity in such a way that the whole manifold is divided into two parts: the interior and the exterior of the black hole. A point is said to be interior if all future geodesics through the point have a non-empty intersection with the singularity. A point is exterior if it is not interior. An important subset of the space is the horizon: the boundary between these two parts.

Most black hole singularities usually emerge from metric considerations: a singularity corresponds to the set of points where a metric invariant diverges. This is however not always the case, as shown for example in the BTZ black holes [5,4]. The latter are obtained as quotients of three-dimensional anti-de Sitter space ( $\mathrm{AdS}_{3}$ ), identified with the (universal covering of the) group manifold $S L_{2}(\mathbb{R})$, under the action of particular subgroups of its isometry group. Regions where the orbits of the identification subgroup are time-like are excluded from the original space

[^0]in order to avoid closed time-like curves. The singularity then corresponds to the region where the identifications become light-like, and it can be shown that the resulting space is indeed a black hole. It was further noticed [11, $8-10$ ] that for a particular subclass of BTZ black holes, the non-rotating massive ones, singularity and horizons have a simple group-theoretical interpretation: they correspond to unions of minimal parabolic (solvable) subgroups of $S L_{2}(\mathbb{R})$, and their translated. It is worth remarking that seeing $S L_{2}(\mathbb{R})$ as the homogeneous space $S O(2,2) / S O(2,1)$, the singularity can be identified with the closed orbits of minimal parabolic subgroups of the isometry group $S O(2,2) \sim\left(S L_{2}(\mathbb{R}) \times S L_{2}(\mathbb{R})\right) / \mathbb{Z}_{2}$. This type of homogeneous space furthermore belongs to the general class of causal symmetric spaces (for a definition and examples, see [12,14,17]).

This observation motivates the following definition:
Definition 1. A causal solvable symmetric black hole is a causal symmetric space where the closed orbits of minimal parabolic subgroups of its isometry group define a black hole singularity.

In this situation, the black hole causal structure is thus completely determined by the action of a solvable group. This observation gives prominence to potential embeddings of these spaces in the framework of noncommutative geometry, in defining noncommutative causal black holes (see also [9]) through the existence of universal deformation formulae for solvable group actions which have been obtained in the context of WKB quantization of symplectic symmetric spaces [6,7].

Non-rotating massive BTZ black holes turn out to enter the class of causal symmetric solvable black holes. The purpose of this paper is to generalize known results about the three-dimensional case to anti-de Sitter spaces of arbitrary dimension, in proving the following theorem:

Theorem 2. For all $l \geq 3$, anti-de Sitter space in $l$ dimensions, seen as the symmetric space $\operatorname{SO}(2, l-1) / S O(1, l-1)$, becomes a causal symmetric solvable black hole, as defined above, when closed orbits of a minimal parabolic subgroup of $\operatorname{SO}(2, l-1)$ and its Cartan conjugate are said to be singular.

This paper intends to prove this theorem. It could be interesting to relate this construction to previous ones giving rise to black holes in AdS spaces; see [13,18,3,2,1,15].

## 2. Symmetric structure

For the sake of notational simplicity, we put $G=S O(2, l-1)$ and $H=S O(1, l-1)$; our space of interest is $\operatorname{AdS}_{l}=G / H$. The equivalence class of $g \in G$ is denoted by $[g]$ and $\vartheta=[e]$. The anti-de Sitter space can be seen as the surface

$$
M \equiv u^{2}+t^{2}-x^{2}-y^{2}-x_{3}^{2}-\cdots-x_{l-1}^{2}=1
$$

embedded in $\mathbb{R}^{2, l-1}$.
When $H$ is the $S O(1, l-1)$ subgroup of $S O(2, l-1)$ which leaves unchanged the vector $(1,0, \ldots, 0) \in \mathbb{R}^{2, l-1}$, the isomorphism is given by

$$
[g] \rightarrow g \cdot\left(\begin{array}{c}
1 \\
0 \\
\vdots \\
0
\end{array}\right)
$$

where the dot means the usual matrix time vector product in $\mathbb{R}^{l+1}$. The Lie algebras of $G$ and $H$ will be denoted by $\mathcal{G}$ and $\mathcal{H}$ respectively. A Cartan involution $\theta: \mathcal{G} \rightarrow \mathcal{G}$ gives rise to a Cartan decomposition

$$
\mathcal{G}=\mathcal{K} \oplus \mathcal{P}
$$

and an involutive automorphism $\sigma=\left.\left.\mathrm{id}\right|_{\mathcal{H}} \oplus(-\mathrm{id})\right|_{\mathcal{Q}}$ to a reductive symmetric space decomposition

$$
\mathcal{G}=\mathcal{H} \oplus \mathcal{Q}
$$

with

$$
[\mathcal{H}, \mathcal{H}] \subset \mathcal{H}, \quad[\mathcal{H}, \mathcal{Q}] \subset \mathcal{Q}, \quad[\mathcal{Q}, \mathcal{Q}] \subset \mathcal{H}
$$

We choose them in such a way that $[\sigma, \theta]=0$.
We now consider Iwasawa decompositions $\mathcal{H}=\mathcal{A}_{\mathcal{H}} \oplus \mathcal{N}_{\mathcal{H}} \oplus \mathcal{K}_{\mathcal{H}}$ and $\mathcal{G}=\mathcal{A} \oplus \mathcal{N} \oplus \mathcal{K}$ taken in such a way that $\mathcal{A}_{\mathcal{H}} \subset \mathcal{A}$ and $\mathcal{N}_{\mathcal{H}} \subset \mathcal{N}$. We denote by $A, N, K, A_{H}, N_{H}$ and $K_{H}$ the analytic connected subgroups of $G$ whose Lie algebras are $\mathcal{A}, \mathcal{N}, \mathcal{K}, \mathcal{A}_{\mathcal{H}}, \mathcal{N}_{\mathcal{H}}$, and $\mathcal{K}_{\mathcal{H}}$ respectively. We denote by $\bar{N}$ the conjugate by $\theta$ of $N$.

The subgroups $R \equiv A N$ and $\bar{R} \equiv A \bar{N}$ are minimal parabolic (solvable) subgroups of $G$, as mentioned in Definition 1. The closed orbits of these two groups define the singularity.

## 3. Causal structure and black hole existence

Definition 3. A light ray is a light-like geodesic.
Lemma 4. Let $E$ be a nilpotent element in $\mathcal{Q}$, and let $\pi: G \rightarrow G / H$ be the canonical projection; then a light ray through $[g] \in \operatorname{AdS}_{l}$ has the form

$$
\begin{equation*}
s(t)=\pi\left(g \mathrm{e}^{-t \mathrm{Ad}(k) E}\right) \tag{1}
\end{equation*}
$$

for a certain $k \in K_{H}$.
Proof. The general form

$$
s(t)=\pi\left(\mathrm{e}^{t X}\right)
$$

is proven in Theorem 3.2 of chapter XI in [16].
Any vector of the form $\operatorname{Ad}(k) E$ has (Killing) zero norm because the trace of a nilpotent matrix vanishes. Now it is easy to see that the only nilpotent matrices in $\mathcal{Q}$ are of the form $q_{0}+q_{i}$ for a certain $i$. $\operatorname{Then} \operatorname{Ad}(k) E=$ $q_{0}+w_{1} q_{1}+\cdots+w_{l-1} q_{l-1}$. From explicit computation given in the Appendix, $\operatorname{Ad}(k) E$ is a generic zero normed vector in $\mathcal{Q}$; see Eq. (A.8).

Points $s(t)$ with $t \in \mathbb{R}^{+}$are said to lie on future-directed light rays issued from $g$.

### 3.1. Search for closed $R$-orbits

Let us start this section by computing the closed orbits of the action of $A N$ and $A \bar{N}$ on $\operatorname{AdS}_{l}$. In order to see if $x=[g] \in M$ lies in a closed orbit of $A N$, we "compare" the basis $\left\{d \pi d L_{g} q_{i}\right\}$ of $T_{x} M$ and the space spanned by the fundamental vectors of the action. If these two spaces are the same, then $x$ belongs to an open orbit (because a submanifold is open if and only if it has the same dimension as the main manifold). This idea is precisely contained in the following proposition.

Proposition 5. If $R$ is a subgroup of $G$ with Lie algebra $\mathcal{R}$, then the orbit $R \cdot \vartheta$ is open in $G / H$ if and only if the projection $\mathrm{pr}: \mathcal{R} \rightarrow \mathcal{Q}$ is surjective.

In order to check the openness of the $R$-orbit of $[g]$, we look at the openness of the $\operatorname{Ad}\left(g^{-1}\right) R$-orbit of $\vartheta$ using the proposition.

A great simplification is possible. The $A N$-orbits are trivially $A N$-invariant. So the $K$ part of $[g]=a n k$ alone fixes the orbit to which $[g]$ belongs. In the explicit parametrization of $K$, we know that the $S O(n)$ part is "killed" by the quotient with respect to $S O(1, n)$. Thus we are left with at most one $A N$-orbit for each element in $S O$ (2). Computations using Proposition 5 show that the closed orbits are given by

$$
\begin{equation*}
\mathscr{S}=\{ \pm[A N], \pm[A \bar{N}]\} . \tag{2}
\end{equation*}
$$

### 3.2. More about the light cone

Lemma 4 claims that if $E$ is nilpotent in $\mathcal{Q}$, then $\{\operatorname{Ad}(k) E\}_{k \in K_{H}}$ is the set of all the light-like vectors in $T_{[\vartheta]} \mathrm{AdS}_{l} \simeq \mathcal{Q}$. So we define the future light cone of $\vartheta$ by

$$
C_{[\vartheta]}^{+}=\left\{\pi\left(\mathrm{e}^{-t \mathrm{Ad}(k) E}\right)\right\}_{\substack{t \in \mathbb{R}^{+} \\ k \in K_{H}}}
$$

and that of a general element $[g] \in \mathrm{AdS}_{l}$ is obtained by the (isometric) action of $g$ thereon:

$$
\begin{equation*}
C_{\pi(g)}^{+}=\left\{\pi\left(g \mathrm{e}^{-t \mathrm{Ad}(k) E}\right)\right\}_{\substack{t \in \mathbb{R}^{+} \\ k \in K_{H}}} . \tag{3}
\end{equation*}
$$

The denomination "future" refers to the fact that it only contains positive $t$. Past light cones correspond to negative $t$. It should be noted that this definition is independent of the choice of the representative $g$ in the class $\pi(g)$ because, for any $h \in H, \pi\left(g h \mathrm{e}^{-t \operatorname{Ad}(k) E}\right)=\pi\left(g h \mathrm{e}^{-t \mathrm{Ad}(k) E} h^{-1}\right)$ which is simply a reparametrization in $K_{H}$.

### 3.3. Computation of the singularity

We here explicitly use the description of $\mathrm{AdS}_{l}$ using the fundamental (defining) representation of $S O(2, n-1)$, i.e. in terms of the embedding coordinates $\left(u, t, x, y, x_{3}, \ldots, x_{l-1}\right) \in \mathbb{R}^{2, l-1}$, and the choices of generators related in the Appendix.

Proposition 6. In terms of the embedding of $\mathrm{AdS}_{l}$ in $\mathbb{R}^{2, l-1}$, the closed orbits of $A N \subset S O(2, l-1)$ are located at $y-t=0$. Similarly, the closed orbits of $A \bar{N}$ correspond to $y+t=0$. In other words, the equation

$$
\begin{equation*}
t^{2}-y^{2}=0 \tag{4}
\end{equation*}
$$

describes the singularity $\mathscr{S}=\mathscr{S}_{A N} \cup \mathscr{S}_{A \bar{N}}$.
Proof. The different fundamental vector fields of the $A N$ action can be computed using $X_{[g]}^{*}=-X g \cdot \vartheta$. For example, in $\mathrm{AdS}_{3}$,

$$
\begin{aligned}
M_{[g]}^{*} & =\left(\begin{array}{cccc}
0 & -1 & 0 & 1 \\
1 & 0 & -1 & 0 \\
0 & -1 & 0 & 1 \\
1 & 0 & -1 & 0
\end{array}\right)\left(\begin{array}{c}
u \\
t \\
x \\
y
\end{array}\right)=\left(\begin{array}{c}
-t+y \\
u-x \\
-t+y \\
u-x
\end{array}\right) \\
& =(y-t) \partial_{u}+(u-x) \partial_{t}+(y-t) \partial_{x}+(u-x) \partial_{y} .
\end{aligned}
$$

The full results are

$$
\begin{align*}
& J_{1}^{*}=-y \partial_{t}-t \partial_{y}  \tag{5a}\\
& J_{2}^{*}=-x \partial_{u}-u \partial_{x}  \tag{5b}\\
& M^{*}=(y-t) \partial_{u}+(u-x) \partial_{t}+(y-t) \partial_{x}+(u-x) \partial_{y}  \tag{5c}\\
& L^{*}=(y-t) \partial_{u}+(u+x) \partial_{t}+(t-y) \partial_{x}+(u+x) \partial_{y}  \tag{5d}\\
& W_{i}^{*}=-x_{i} \partial_{t}-x_{i} \partial_{y}+(y-t) \partial_{i}  \tag{5e}\\
& V_{j}^{*}=-x_{j} \partial_{u}-x_{j} \partial_{x}+(x-u) \partial_{j}, \quad i, j=3, \ldots, l-1 . \tag{5f}
\end{align*}
$$

First consider points satisfying $t-y=0$. It is clear that, at these points, the $l$ vectors $J_{1}^{*}, M^{*}, L^{*}$ and $W_{i}^{*}$ are linearly dependent. Then, there are at most $l-1$ linearly independent vectors amongst the $2(l-1)$ vectors (5); thus the points belong to a closed orbit.

We now show that a point with $t-y \neq 0$ belongs to an open orbit of $A N$. It is easy to see that $J_{1}^{*}, L^{*}$ and $M^{*}$ are three linearly independent vectors. The vectors $V_{i}^{*}$ give us $l-3$ more. Then they span an $l$-dimensional space.

The same can be done with the closed orbits of $A \bar{N}$. The result is that a point belongs to a closed orbit of $A \bar{N}$ if and only if $t+y=0$.

For the three-dimensional case, it was shown in $[4,11]$ that the non-rotating BTZ black hole singularity is precisely given by Eq. (4). Hence, the following is a particular case of Theorem 2:

Corollary 7. The non-rotating BTZ black hole is a causal symmetric solvable black hole.

### 3.4. Existence of a horizon

We first consider points of the form $K \cdot \vartheta$, which are parametrized by an angle $\mu$. Up to the choice of this parametrization, a light-like geodesic through $\mu$ is given by

$$
\begin{equation*}
K \cdot \mathrm{e}^{-s \operatorname{Ad}(k) E} \cdot \vartheta \tag{6}
\end{equation*}
$$

with $k \in S O(l-1)$ and $s \in \mathbb{R}$.
This geodesic reaches $\mathscr{S}_{A N}$ and $\mathscr{S}_{A \bar{N}}$ for values $s_{A N}$ and $s_{A \bar{N}}$ of the affine parameter, given by

$$
\begin{equation*}
s_{A N}=\frac{\sin \mu}{\cos \mu-\cos \alpha}, \quad \text { and } \quad s_{A \bar{N}}=\frac{\sin \mu}{\cos \mu+\cos \alpha} \tag{7}
\end{equation*}
$$

where $\cos \alpha$ is the $w_{2}$ element of the matrix in (A.8) $\left(-1 \leq w_{2} \leq 1\right)$.
Because the part $\sin \mu=0$ is $\mathscr{S}_{A N}$, we may restrict ourselves to the open connected domain of $\operatorname{AdS}_{l}$ given by $\sin \mu>0$. More precisely, $\sin \mu=0$ is the equation of $\mathscr{S}_{A N}$ in the $A N K$ decomposition. In the same way, $\mathscr{S}_{A \bar{N}}$ is given by $\sin \mu^{\prime}=0$ in the $A \bar{N} K$ decomposition. In order to escape the singularity, the point $\mu$ needs $s_{A N}, s_{A \bar{N}}<0$. It is only possible to find directions (i.e. an angle $\alpha$ ) which respect this condition when $\cos \mu<0$. So the point $\cos \mu=0$ is one point of the horizon.

### 3.5. A characterization of the horizon

Let $D[g]$ be the set of light-like directions (vectors in $S O(n)$ ) for which the point $[g]$ falls into $\mathscr{S}_{A N}$. Similarly, $\bar{D}[g]$ is the set of directions for which it falls into $\mathscr{S}_{A \bar{N}}$. It is actually possible to express $\bar{D}$ in terms of $D$. Indeed

$$
\begin{array}{ll}
k \in \bar{D}[g] & \text { iff } \pi\left(g \mathrm{~g}^{t k \cdot E}\right) \in \mathscr{S}_{A \bar{N}} \\
& \text { iff } \pi\left(\theta(g) \theta\left(\mathrm{e}^{\left.\left.t \cdot k \cdot E_{1}\right)\right)}\right) \in \mathscr{S}_{A N}\right.  \tag{8}\\
\text { iff } & \theta(k) \in D(\theta[g]) \\
\text { iff } & k \in(D(\theta[g]))_{\theta} .
\end{array}
$$

So

$$
\begin{equation*}
\bar{D}[g]=(D \theta[g])_{\theta} \tag{9}
\end{equation*}
$$

where the definition of $k_{\theta}$ is

$$
\theta(\operatorname{Ad}(k) E)=\operatorname{Ad}\left(k_{\theta}\right) E
$$

This definition is possible because $\theta$ is an inner automorphism.
It is easy to see that $\theta$ changes the sign of the spatial part of $k$, i.e. changes $w_{i} \rightarrow-w_{i}$.
How do we express the condition $g \in \mathscr{H}$ in terms of $D[g]$ ? The condition for being in the black hole is $D[g] \cup$ $\bar{D}[g]=S O(n)$. If the complementary of $D[g] \cup \bar{D}[g]$ has an interior (i.e. if it contains an open subset), then by continuity the complementary $D\left[g^{\prime}\right] \cup \bar{D}\left[g^{\prime}\right]$ has also an interior for all $\left[g^{\prime}\right]$ near $[g]$. In this case, $[g]$ cannot belong to the horizon. So a characterization of $\mathscr{H}$ is the fact that the boundary of $D[g]$ and $\bar{D}[g]$ coincide. Eq. (9) shows that $\mathscr{H}$ is $\theta$-invariant.

We can explicitly express $D[\mu]$ for $\mu \in S O(2)$ by examining Eq. (7). Let us write $w_{2}$ instead of $\cos \alpha$. The set $D[\mu]$ is the set of $w_{2} \in[-1,1]$ such that $\cos \mu-w_{2}>0$ :

$$
\begin{equation*}
D[\mu]=[-1, \cos \mu[. \tag{10}
\end{equation*}
$$

So in order for $[\mu]$ to belong to $\mathscr{H}$, it must satisfy

$$
D[\theta]_{\theta}=\left[-1, \cos \mu^{\prime}\left[_{\theta}=\right]-\cos \mu^{\prime}, 1\right] .
$$

Consequently, if $\mu$ is the $K$ component of $g$ in the ANK decomposition and $\mu^{\prime}$ that of $\theta u$, then we can describe the horizon by

$$
\begin{equation*}
\cos \mu=-\cos \mu^{\prime} \tag{11}
\end{equation*}
$$

## Acknowledgments

We are grateful to Pierre Bieliavsky for suggesting the problem. We also thank him, as well as Philippe Spindel, for numerous enlightening discussions. We would also like to thank the "Service de Physique Théorique" of the Université Libre de Bruxelles, where part of this work was carried out, for its hospitality.

## Appendix. Explicit matrix choices

The first choice is to parametrize $S O(2, n)$ and $S O(1, n)$ in such a way that the latter leaves unchanged the vector $(1,0,0, \ldots)$. Then

$$
\mathcal{H}=\mathfrak{s o}(1, n) \rightsquigarrow\left(\begin{array}{cc}
0 & 0  \tag{A.1}\\
0 & 0 \\
\left(\begin{array}{cc}
\vdots & \uparrow \\
0 & v \\
\vdots & \downarrow
\end{array}\right) & \binom{\cdots 0 \cdots}{\leftarrow v^{t} \rightarrow} \\
B
\end{array}\right)
$$

where $v$ is $n \times 1$ and $B$ is skew symmetric $n \times n$. When we speak about $\mathfrak{s o}(n)$, we usually refer to the $B$ part of $\mathcal{H}$. A complementary space $\mathcal{Q}$ such that $[\mathcal{H}, \mathcal{Q}] \subset \mathcal{Q}$ is given by

$$
\mathcal{Q} \rightsquigarrow\left(\begin{array}{cc}
0 & a  \tag{A.2}\\
-a & 0 \\
\left(\begin{array}{cc}
\uparrow & \vdots \\
w & 0 \\
\downarrow & \vdots
\end{array}\right) & \\
\cdots 0 \cdots
\end{array}\right)
$$

We consider the involutive automorphism $\sigma=\operatorname{id}_{\mathcal{H}} \oplus(-\mathrm{id})_{\mathcal{Q}}$ and the corresponding symmetric space structure on $\mathcal{G}$. As a basis of $\mathcal{Q}$, we choose $q_{0}$ as the $2 \times 2$ antisymmetric upper left square and $q_{i}$ as those obtained with $w$ full of zeros apart from a 1 as the $i$ th component. Next we choice the Cartan involution $\theta(X)=-X^{t}$ which gives rise to a Cartan decomposition

$$
\mathcal{G}=\mathcal{K} \oplus \mathcal{P} .
$$

The latter choice is made in such a way that $[\sigma, \theta]=0$. It can be computed, but it is not astonishing that the compact part $\mathcal{K}$ is made up of "true" rotations while $\mathcal{P}$ contains the boost. So

$$
\mathcal{K}=\left(\begin{array}{cc}
\mathfrak{s o}(2) & \\
& \mathfrak{s o}(n)
\end{array}\right)
$$

where elements of $S O(2)$ are represented as

$$
\left(\begin{array}{cc}
\cos \mu & \sin \mu \\
-\sin \mu & \cos \mu
\end{array}\right)
$$

A common abuse of notation in the text is to identify the angle $\mu$ with the element of $S O(2)$ itself.
In order to build an Iwasawa decomposition, one has to choose a maximal abelian subalgebra $\mathcal{A}$ of $\mathcal{P}$. Since rotations are in $\mathcal{K}$, they must be boosts, and the fact that there are only two time-like directions restricts $\mathcal{A}$ to a twodimensional algebra. Up to reparametrization, it is thus generated by $u \partial_{x}+x \partial_{u}$ and $t \partial_{y}+y \partial_{t}$. Our matrix choices
are

$$
J_{1}=\left(\begin{array}{cccc} 
& 0 & & \\
0 & 0 & 0 & 1 \\
& 0 & & \\
& 1 & &
\end{array}\right) \in \mathcal{H}, \quad \text { and } \quad J_{2}=q_{1}=\left(\begin{array}{llll}
0 & 0 & 1 & 0 \\
0 & & & \\
1 & & & \\
0 & & &
\end{array}\right) \in \mathcal{Q}
$$

From here, we have to build root spaces. There still remain a lot of arbitrary choices - among them, the positivity notion on the dual space $\mathcal{A}^{*}$. An element $X$ in $\mathcal{G}_{(a, b)}$ fulfills $\operatorname{ad}(X) J_{1}=a J_{1}$ and $\operatorname{ad}(X) J_{2}=b J_{2}$. The symbol $E_{i j}$ denotes the matrix full of zeros with a 1 as the component $i j$. The results are

$$
\mathcal{G}_{(0,0)} \rightsquigarrow\left(\begin{array}{lllll} 
& & x & 0 &  \tag{A.3}\\
& & 0 & y & \\
x & 0 & & & \\
0 & y & & & \\
& & & & D
\end{array}\right),
$$

where $D \in M_{(n-2) \times(n-2)}$ is skew symmetric,

$$
\begin{align*}
& \mathcal{G}_{(1,0)} \rightsquigarrow W_{i}=E_{2 i}+E_{4 i}+E_{i 2}-E_{i 4},  \tag{A.4a}\\
& \mathcal{G}_{(-1,0)} \rightsquigarrow Y_{i}=-E_{2 i}+E_{4 i}-E_{i 2}-E_{i 4},  \tag{A.4b}\\
& \mathcal{G}_{(0,1)} \rightsquigarrow V_{i}=E_{1 i}+E_{3 i}+E_{i 1}-E_{i 3},  \tag{A.4c}\\
& \mathcal{G}_{(0,-1)} \rightsquigarrow X_{i}=-E_{1 i}+E_{3 i}-E_{i 1}-E_{i 3} \tag{A.4d}
\end{align*}
$$

with $i: 5 \rightarrow n+2$ and

$$
\begin{array}{ll}
\mathcal{G}_{(1,1)} \rightsquigarrow M=\left(\begin{array}{cccc}
0 & 1 & 0 & -1 \\
-1 & 0 & 1 & 0 \\
0 & 1 & 0 & -1 \\
-1 & 0 & 1 & 0
\end{array}\right), & \mathcal{G}_{(1,-1)} \rightsquigarrow L=\left(\begin{array}{cccc}
0 & 1 & 0 & -1 \\
-1 & 0 & -1 & 0 \\
0 & -1 & 0 & 1 \\
-1 & 0 & -1 & 0
\end{array}\right), \\
\mathcal{G}_{(-1,1)} \rightsquigarrow N=\left(\begin{array}{cccc}
0 & 1 & 0 & 1 \\
-1 & 0 & 1 & 0 \\
0 & 1 & 0 & 1 \\
1 & 0 & -1 & 0
\end{array}\right), & \mathcal{G}_{(-1,-1)} \rightsquigarrow F=\left(\begin{array}{cccc}
0 & 1 & 0 & 1 \\
-1 & 0 & -1 & 0 \\
0 & -1 & 0 & -1 \\
1 & 0 & 1 & 0
\end{array}\right) . \tag{A.6}
\end{array}
$$

The choice of positivity is

$$
\begin{equation*}
\mathcal{N}=\left\{V_{i}, W_{j}, M, L\right\} \tag{A.7}
\end{equation*}
$$

The following result is important in the computation of the light cones: if $k \in S O(n)$, then the choice $E=q_{0}+q_{2}$ of the nilpotent element in $\mathcal{Q}$ gives

$$
\operatorname{Ad}(k) E=\left(\begin{array}{ccccc}
0 & 1 & w_{1} & w_{2} & \cdots  \tag{A.8}\\
-1 & & & & \\
w_{1} & & & & \\
w_{2} & & & & \\
\vdots & & & &
\end{array}\right)
$$

where the elements $w_{i}$ are restricted by the condition $\sum_{i=1}^{l-1} w_{i}^{2}=1$.

## References

[1] S. Aminneborg, I. Bengtsson, S. Holst, P. Peldan, Making anti-de sitter black holes, Classical Quantum Gravity 13 (1996) 2707-2714. gr-qc/ 9604005.
[2] M. Banados, Constant curvature black holes, Phys. Rev. D 57 (1998) 1068-1072. gr-qc/9703040.
[3] M. Banados, A. Gomberoff, C. Martinez, Anti-de sitter space and black holes, Classical Quantum Gravity 15 (1998) 3575-3598. hep-th/ 9805087.
[4] M. Banados, M. Henneaux, C. Teitelboim, J. Zanelli, Geometry of the (2+1) black hole, Phys. Rev. D 48 (1993) 1506-1525. gr-qc/9302012.
[5] M. Bañados, C. Teitelboim, J. Zanelli, The black hole in three-dimensional space-time, Phys. Rev. Lett. 69 (1992) 1849-1851. hep-th/ 9204099.
[6] P. Bieliavsky, Strict quantization of solvable symmetric spaces, J. Symplectic Geom. 1 (2) (2002) 269. math.QA/0010004.
[7] P. Bieliavsky, M. Massar, Strict deformation quantization for actions of a class of symplectic lie groups, Progr. Theoret. Phys. Suppl. 144 (2001) 1-21. math.QA/0011144.
[8] P. Bieliavsky, S. Detournay, M. Rooman, P. Spindel, Noncommutative locally anti-de sitter black holes, in: Y. Maeda, N. Tose, N. Miyazaki, S. Watamura and D. Sternheimer (Eds.), Noncommutative Geometry And Physics (Proceedings of the COE International Workshop Yokohama, Japan 26-28 February, 1-3 March 2004), 2005, World Scientific. math.qa/0507157.
[9] P. Bieliavsky, S. Detournay, M. Rooman, P. Spindel, Star products on extended massive non-rotating BTZ black holes, J. High Energy Phys. 06 (2004) 031. hep-th/0403257.
[10] P. Bieliavsky, S. Detournay, M. Rooman, P. Spindel, BTZ black holes, WZW models and noncommutative geometry, in: Proceedings of Rencontres Mathematiques de Glanon, Glanon, France, 5-9 July 2004. e-Print Archive: hep-th/0511080.
[11] P. Bieliavsky, M. Rooman, P. Spindel, Regular Poisson structures on massive non-rotating BTZ black holes, Nuclear Phys. B 645 (2002) 349-364. hep-th/0206189.
[12] J. Faraut, G. Olafsson, Causal semisimple symmetric spaces: The geometry and harmonic analysis, in: K.H. Hofmann, J.D. Lawson, E.B. Vinberg (Eds.), Semigroups in Algebra, Geometry and Analysis, 1995, pp. 3-32.
[13] J. Figueroa-O'Farrill, O. Madden, S.F. Ross, J. Simon, Quotients of ads $(p+1) x s^{* *} q$ : Causally well-behaved spaces and black holes, Phys. Rev. D 69 (2004) 124026. hep-th/0402094.
[14] J. Hilgert, G. Olafsson, Causal Symmetric Spaces, Geometry and Harmonic Analysis, Perspectives in Mathematics 18, Academic Press, 1997.
[15] S. Holst, P. Peldan, Black holes and causal structure in anti-de sitter isometric spacetimes, Classical Quantum Gravity 14 (1997) $3433-3452$. gr-qc/9705067.
[16] S. Kobayashi, K. Nomizu, Foundation of Differential Geometry, vol. 2, Interscience publishers, 1969.
[17] K. Koufany, Analyse et géométrie des espaces symétriques de type Cayley, http://www.iecn.u-nancy.fr/~koufany/archives/cayley_space.ps.
[18] O. Madden, S.F. Ross, Quotients of anti-de sitter space, Phys. Rev. D 70 (2004) 026002. hep-th/0401205.


[^0]:    * Corresponding author.

    E-mail address: claessens@math.ucl.ac.be (L. Claessens).
    1 "Chercheur FRIA", Belgium.

